

Configuration Manual

MSc Research Project Data Analytics

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Configuration Manual

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1 Introduction

The purpose of this document is go through the software, hardware and system configurations which are required during the entire process flow of the research project. Below is the process which is followed to apply pre processing techniques followed by machine and deep learning algorithms.

2 System Configuration

2.1 Hardware

Processor: Intel(R) Core(TM) i5-8300H CPU @ 2.30GHz GPU: NVIDIA GeForce GTX 1050Ti (4GB) RAM: 8GB Storage: 1TB HDD Operating system: Windows 10, 64-bit.

2.2 Software

Python using Google Collaboratory and Jupyter Notebook: Data cleaning, data preprocessing, analysis and visualization.

Microsoft Excel: Save the data from extracted FITS files

Device name	LAPTOP-1B5C81HL
Processor	Intel(R) Core(TM) i5-1035G1 CPU @ 1.00GHz 1.19 GHz
Installed RAM	8.00 GB (7.70 GB usable)
Device ID	1FF55988-2643-426C-8A61-F8E46B2F087B
Product ID	00327-35885-16033-AAOEM
System type	64-bit operating system, x64-based processor
Pen and touch	No pen or touch input is available for this display

Figure 1: System Configuration

3 Implementation

The implementation of the thesis follows a step by step approach as mentioned in Figure 2. The research is divided in three parts. Extracting the data then cleansing it for any null values followed by the next step of applying pre processing techniques of Fast Fourier Transform (FFT) and Recurrence Plot (RP). Once these are applied, machine learning algorithm, Support Vector Machine(SVM) is applied to FFT pre processed data and Convolutional Neural Network along with pre trained VGG 16 model is applied to data pre processed by RP.



Figure 2: End to end flow of finding Exoplanets

4 Data Pre-processing

4.1 Data Collection from Mikulski Archive for Space Telescope (MAST) Portal

The data chosen for this research is from MAST portal 1 where there are different parameters for searching the data. The data is filtered on taking Exoplanet Host star data and False positive data. Figure 3 shows the different filters which can be applied to the data.

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MAST STSct Tools • Mession Search • Kepler Home Mission Getting Started Data S	Search Website 🚺 🖌 Po iearch Target Search 8	allow Us • Register Reliar17 CasJobs	FFI+ Search
Archive Status Kepler Da	ita Search &	Retrieval	Help Field Descriptions
Standard Form		File Upload	Form
Search	Reset	Cle	ar Form
Target Name Right Ascension	Resolv Resolve Declin	er I ivi	Radius.(arcmin) 0.02 Equinox J2000 ♥
Kepler ID	Investigation ID		2Mass.ID
KEP Mag	Target_Type ng Cadence C Short Ca	adence	Release Date
Tetf	Log_Q		Quarter
	Condition Flag Al Targets V		
User-specified field 1 Field Descr	iptions Use	r-specified field 2	Field Descriptions

Figure 3: MAST Portal for NASA's Kepler Mission's Data

¹Kepler Mission Data: http://archive.stsci.edu/kepler/data_search/search.php

4.2 FITS file

The data is in the format of URLs for the light curves. These can be accessed by astropy.io package in Python which extracts the time series data. Figure 4 is a sample of the FITS data.(Silva et al.; 2013)

https://archive.stsci.edu/missions/kepler/lightcurves/0008/0008927	772/kplr000892772-2010078095331_llc.fits
https://archive.stsci.edu/missions/kepler/lightcurves/0008/0008927	772/kplr000892772-2010174085026_11c.fits
https://archive.stsci.edu/missions/kepler/lightcurves/0008/0008927	772/kplr000892772-2010265121752_llc.fits
https://archive.stsci.edu/missions/kepler/lightcurves/0008/0008927	772/kplr000892772-2010355172524 llc.fits
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https://archive.stsci.edu/missions/kepler/lightcurves/0008/0008927	772/kplr000892772-2011177032512 llc.fits
https://archive.stsci.edu/missions/kepler/lightcurves/0008/0008927	772/kplr000892772-2011271113734 llc.fits
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https://archive.stsci.edu/missions/kepler/lightcurves/0008/0008927	772/kp1r000892772-2012088054726 11c.fits
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https://archive.stsci.edu/missions/kepler/lightcurves/0008/0008927	772/kplr000892772-2012277125453 llc.fits
https://archive.stsci.edu/missions/kepler/lightcurves/0008/0008927	772/kp1r000892772-2013011073258 11c.fits
https://archive.stsci.edu/missions/kepler/lightcurves/0008/0008927	772/kp1r000892772-2013098041711 llc.fits
https://archive.stsci.edu/missions/kepler/lightcurves/0008/0008927	772/kp1r000892772-2013131215648 11c.fits
https://archive.stsci.edu/missions/kepler/lightcurves/0010/0010259	986/kp1r001025986-2009131105131 11c.fits
https://archive.stsci.edu/missions/kepler/lightcurves/0010/0010259	986/kp1r001025986-2009166043257 11c.fits
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https://archive.stsci.edu/missions/kepler/lightcurves/0010/0010259	986/kplr001025986=2009350155506_llc_fits
https://archive.stsci.edu/missions/kepler/lightcurves/0010/0010255	986/kplr001025986-2010078095331 llc fits
https://archive.stsci.edu/missions/kepier/lightcurves/0010/0010255	986/kp1r001025986=2010174085026_11c_fite
https://archive.stsci.edu/missions/kepier/lightcurves/0010/0010255	006/kp1r001025006=2010265121752_llc_fite
https://archive.staci.edu/missions/kepier/lightsurves/0010/0010255	006/kpir001025986 2010265121752_110.1105
https://archive.stsci.edu/missions/kepier/lightcurves/0010/0010255	006/lml=001025500 20105551/2524_110.1105
nucps.//archive.susci.edu/missions/kepier/lightcurves/0010/0010259	200/ KDITOOTOZ 2000-20110/3133259_11C.11C3

Figure 4: FITS file having URLs

To do processing of data using Python libraries like astropy, Scipy, Keras, Tensorflow, CV2, etc need to be installed. These need to be imported and then used. The interface used for Python in this research id Google Collaboratory. The notebook can also be run on Anaconda's Jupyter Notebook. In below images Figure 5 and Figure 6 all libraries required for the project are imported.



Figure 5: Python Libraries Imported

To access these files these are read using the fits.open() and the brightness of the star which is signified by PDCSAP_FLUX feature. The NAN values are replaced by the mean flux values as below in Figure 7.



Figure 6: Python Libraries Imported

<pre># # The data is read from the links and stored in variable called Data_Exo and Data_No_Exo # # This data is in FITS(Flexible Image Transport System) format and can be accessed using astropy lib in python.</pre>
This library is imported in the first cell.
<pre># # Saving data from the links to CSV file in time series # # Total data contains number of data sets used for testing and training.</pre>
Training data is 80% of Total data sat whereas Tacting data is reamining 20%
Enter split ratio for training data
<pre># POS = pd.DataFrame()</pre>
<pre># for i in range(DATA_TO_COPY2CSV):</pre>
<pre># Data_Exo = fits.open(Exo_content_list[i])</pre>
<pre># flux = Data_Exo[1].data[' '].byteswap().newbyteorder()</pre>
<pre># mean = np.nanmean(flux)</pre>
<pre># inds = np.where(np.isnan(flux))</pre>
<pre># flux[inds] = mean</pre>
<pre># POS = POS.append(pd.DataFrame(flux.reshape(1,flux.shape[0]), dtype='float32'))</pre>
<pre># print("DONE ",i+1,"/",str(DATA_TO_COPY2CSV))</pre>
<pre># POS = POS.fillna(0)</pre>
<pre># POS.to_csv('POS_TimeSeries.csv', mode='w', header=False, index=False)</pre>
<pre># print("DONE POS")</pre>
POS
NEG

Figure 7: Cleansing data and coverting it to CSV format

4.3 Plot for Time Series Data

Using the time and flux values the initial data is plotted to check how patterns are being formed. For experiment 1, data with initial cleansing was applied to SVM. The results were satisfactory with a classification accuracy of 52 percent. To improve this accuracy this trsearch focuses on using pre processing techniques using FFT and RP.



Figure 8: Applying SVM without any pre processing

4.4 Application of Fast Fourier Transform

The next experiment uses initially cleansed data. FFT is applied to this data by using Scipy library in Python which provides fft package. Using fftpack and fftfreq, FFT is applied in samples. A sample of 100 data points is created and iterated for all the data.

```
# # FFT analysis
X_train_fft = np.abs(fft(X_train)).astype(np.float32)
X_test_fft = np.abs(fft(X_test)).astype(np.float32)
def Plot(X, i):
 f s = 100
 freqs = fftpack.fftfreq(len(X_train_fft[i])) * f_s
 plt.xlabel('Frequency in Hertz [Hz]')
 plt.ylabel('Frequency Domain (Spectrum) Magnitude')
 # ax.set xlim(-f_s / 2, f_s / 2)
 plt.ylim(0, 10000)
 plt.plot(freqs, X_train_fft[i], marker=".", markersize=2)
 plt.figure()
Plot(X_train_fft[10],10)
Plot(X_train_fft[0],0)
Plot(X_train_fft[30],30)
Plot(X train fft[-11],-11)
Plot(X train fft[-21],-21)
Plot(X train fft[-31],-31)
```

Figure 9: FFT applied to Samples of Data

4.5 Application of Recurrence Plots

Recurrence Plots plot time series data and help in visualizing the data. To analyse the data it was first converted to 1 dimensional figure which gave different patterns of light

curves. Using matplotlib and CV2 libraries in Python recurrence plots are plotted as shown in Figure 10.

```
# Reference Plots from time curves
# 1, 2, 3 recurrence plots are for postive data and 4, 5, 6 are for negative data
def rec_plot(s, eps=0.50, steps=10):
   d = pdist(s[:,None])
   d = np.floor(d/eps)
   d[d>steps] = steps
   Z = squareform(d)
   return Z
List = [10, 0, 30]
for _, i in enumerate(List):
 image = rec_plot(X_train[i,:2000])
 plt.figure()
 plt.imshow(image,cmap="gray")
List = [X_train.shape[0]-11, X_train.shape[0]-21, X_train.shape[0]-31]
for _, i in enumerate(List):
  image = rec_plot(X_train[i,:2000])
 plt.figure()
 plt.imshow(image,cmap="gray")
```

Figure 10: Recurrence Plot applied for 1D image

These images cannot be fed to any model or algorithm as they are in 1 Dimensional format. To make these images ready for model application it is needed that they are converted to 2 dimensional or 3 dimensional format. This is done by using stacking technique present in CV2 library, The interpolation of images are changed and made suitable for modelling.

```
# Stacking recurrence plot images (Conversion from single channel to multi channel images)
# 3 Channel image for Positive data
image_pos = rec_plot(X_train[0,:3000])
print("Shape of image before resizing",image_pos.shape)
image_pos = cv2.resize(image_pos, (HEIGHT,WIDTH), interpolation=cv2.INTER_CUBIC)
print("Shape of image before resizing",image_pos.shape)
image_pos = np.stack((image_pos, image_pos, image_pos),axis=0).reshape(HEIGHT,WIDTH,3)
plt.figure()
plt.imshow(image_pos)
# 3 Channel image for Negative data
image_neg = rec_plot(X_train[-11,:3000])
print("Shape of image before resizing",image_neg.shape)
image_neg = cv2.resize(image_neg, (HEIGHT,WIDTH), interpolation=cv2.INTER_CUBIC)
print("Shape of image before resizing",image_neg.shape)
image_neg = np.stack((image_neg, image_neg, image_neg),axis=0).reshape(HEIGHT,WIDTH,3)
plt.figure()
plt.imshow(image_neg)
```

Figure 11: Converting RP to 2 dimensional images

After this SVM is applied to FFT processed data and CNN is applied to RP processed data. CNN uses a pre-trained model VGG16 with customized layers for classification of the data. All three steps were performed using Keras and Tensorflow libraries. Following

<pre>base_model = VGG16(weights="imagenet", include_top=False, input_shape=X_TRAIN[0].shape)</pre>
<pre>base_model.trainable = False ## Not trainable weights</pre>
<pre>## Preprocessing input train_ds = preprocess_input(X_TRAIN) test_ds = preprocess_input(X_TEST)</pre>
Downloading data from https://storage.googleapis.com/tensorflow/keras-applications/vgg16/vgg16_weights_tf_dim_ordering_tf_kernels_notop.h5 58892288/55889256 [
Loading VGG16 model base_model = VGG16(weights=" <mark>imagenet</mark> ", include_top=False, input_shape=(HEIGHT, WIDTH, 3)) base_model.trainable = False ## Not trainable weights
<pre>## Preprocessing input # train_ds = preprocess_input(X_TRAIN.reshape(X_TRAIN.shape[0],HEIGHT,WIDTH,3)) # test_ds = preprocess_input(X_TEST.reshape(X_TEST.shape[0],HEIGHT,WIDTH,3))</pre>

Figure 12: Create base model and Pre process input for VGG16

figures are how model is loaded for VGG16 followed by application of dense layers and model compliation by using Adam optimizer and crossentropy function.

The snippet in Figure 12 executes the base model with pre pressing for the model. It categorises the data into positive and negative classes.



Figure 13: Creating and applying dense layers to the model

The next code snippet in Figure 13 executes customized layers that are built for our model four dense layers and reLU activation function. A prediction layer is also used which uses the Softmax activation function. The total number of parameters are found and segregated into trainable and non trainable parameters.

<pre>model.compile(optimizer-'Adam', loss-'categorical_crossentropy', metrics=['accuracy'],)</pre>	
<pre>es = EarlyStopping(monitor='val_accuracy', mode='max', patience=5, restore_best</pre>	_weights=True)
<pre>history2 = model.fit(train_ds, train_labels, epochs=50, validation_split=0.2, ba</pre>	tch_size=32, callbacks=[es])

Figure 14: Model compilation

Once the parameters are found, Figure 14 optimizes the model by using optimizer, cross entropy function and loss function.

The FFT processed data is applied with SVM again. The snippet below is applying SVM on the data which is processed by FFT previously.

<pre>def model(classifier,dtrain_x,dtrain_y,dtest_x,dtest_y): #fit the model classifier.fit(dtrain_x,dtrain_y) predictions = classifier.decision_function(dtest_x) predict_acc = classifier.predict(dtest_x)</pre>
<pre>#Accuracy print ("\naccuracy_score :",accuracy_score(dtest_y,predict_acc))</pre>
<pre>#Classification report print ("\nclassification report :\n",(classification_report(dtest_y,predict_acc)))</pre>
<pre>#Confusion matrix plt.figure(figsize(13,10)) plt.subplt(21) sns.heatmap(confusion_matrix(dtest_y,predict_acc),annot=True,cmap="viridis",fmt = "d",linecolor="k",linewidths=3) plt.title"(CONFUSION MATRIX",fontsize=20) return predictions</pre>
<pre>SVM_model=SVC() predictions3 = model(SVM_model,</pre>
<pre>X_test_fft.reshape(X_test_fft.shape[0]),X_test_fft.shape[1]), Y_test_fft.reshape(Y_test_fft.shape[0]))</pre>

Figure 15: Model compilation

5 Evaluation

Once the models are built, it was compared with normal model and pre processed models. Evaluation of models was done on the basis accuracy, precision and ROC curve. Confusion matrix is also used as a parameter. Below snippet is used to create ROC curve for all the models and was used as a metrics to measure the performance.





Figure 16: Code for ROC curve for RP

Figure 17: ROC Curve for RP

ROC curve is also created for before application of FFT and after application of FFT. This curve was created together in one graph to compare in a better way.

Using above codes ROC curves were created and in a similar way Confusion matrix and other metrics were calculated for each model.





Figure 18: Code for ROC curve for FFT Figure 19: Combined ROC for SVM before and after FFT

References

Silva, D. F., Souza, V. M. D. and Batista, G. E. (2013). Time Series Classification Using Compression Distance of Recurrence Plots, 2013 IEEE 13th International Conference on Data Mining, pp. 687–696. ISSN: 2374-8486.